

Nuclear structure functions

TUS K. Saito

- DIS kinematics what can we see in DIS ?
- Experiments what is the nuclear EMC effect ?
- Theoretical approaches can we understand it ?
- Possible explanations what do we need ?
- Summary



1. Kinematics of Deep Inelastic Scattering (DIS)





Initial and final lepton 4-momentum:

$$k^{\mu}, k'^{\mu}, k^{2} = k'^{2} = m^{2} \approx 0$$

- •Virtual photon 4-momentum squared: $q^2 = (k - k')^2 \equiv -Q^2 < 0$
- •Initial nucleon (nucleus) 4-momentum: $P^{\mu} = (E_T, \vec{P}), P^2 = M_T^2$
- Final hadronic 4-momenyum squared: $P_X^2 = (P + q)^2 \equiv W^2$
- •Inelasticity (energy transfer in Lab): $v = (P \cdot q) / M_T$
- •Bjorken variable: $0 < x = Q^2 / 2M_T v \le 1$

KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 2/19

• The differential cross section (unpolarized):

$$\frac{d^{2}\sigma}{d\Omega dE'} = \frac{\alpha^{2}E'}{Q^{4}E} L_{\mu\nu} W^{\mu\nu}$$

lepton tensor (symmetric part):

$$L^{\mu\nu} = 2(k^{\mu}k'^{\nu} + k'^{\mu}k^{\nu} - k \cdot k'g^{\mu\nu})$$

hadronic tensor (symmetric part):

$$W^{\mu\nu} = W_1(x,Q^2)e^{\mu\nu} + [W_2(x,Q^2)/M_T^2]\widetilde{P}^{\mu}\widetilde{P}^{\nu},$$

$$(e^{\mu\nu} = g^{\mu\nu} - q^{\mu}q^{\nu}/q^2, \quad \widetilde{P}^{\mu} = P^{\mu} - (P \cdot q)q^{\mu}/q^2)$$

• Structure functions F1 and F2:

$$F_1(x,Q^2) = M_T W_1(x,Q^2), \ F_2(x,Q^2) = (P \cdot q / M_T) W_2(x,Q^2)$$

• Bjorken limit:

$$F_{1,2}(x,Q^2) \xrightarrow{Q^2 \to \infty, v \to \infty, x: fixed} \to F_{1,2}(x)$$
 Bjorken scaling







• What can we see in DIS ?



The approximate Q^2-independence of the structure functions

 \rightarrow the virtual photon sees point-like constituents in the target – quarks

 \rightarrow using distributions of quarks and anti-quarks,

$$F_1(x) = \frac{1}{2} \sum_f e_f^2 [q_f(x) + \overline{q}_f(x)], F_2(x) = 2xF_1(x) \quad \text{(Callan-Gross relation)}$$

The small scaling violation is calculated by pQCD.

DIS probes a current-current correlation in the target ground state. In the Bjorken limit, the probed correlation is light-like:

$$y^{\pm} = (t \pm y_3) / \sqrt{2}, y^- \to 0, \vec{y}_{\perp} \to 0, y^+ \approx \sqrt{2} / M_T x$$

 $|t|, |y^3| \le 0.2(fm) / x \approx \ell_c$

~ 2.0(fm) for x ~ 0.1 ~ 1.0(fm) for x ~ 0.2 ~ 0.4(fm) for x ~ 0.5 ~ 0.2(fm) for x ~ 1.0





KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 4/19

2. Experiments





KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 5/19



KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 6/19





SLAC



KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 7/19







KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 8/19

KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 9/19

3. Theoretical approaches

- 3-1. Effect of the conventional nuclear physics Binding and Fermi motion
- 3-2. Shadowing effect at small x
- 3-3. Anti-shadowing?

3-1. Effect of the conventional nuclear physics — Binding and Fermi motion

How does the conventional nuclear physics affect F2(x)?

The light-cone momentum distribution of N in A: $D_{j(=p,n)/A}^{\alpha}(y, p^{2}) = y \int \frac{d^{4}p'}{(2\pi)^{4}} S_{j}^{\alpha}(p') \delta\left(y - \frac{p' \cdot q}{P_{A} \cdot q} \frac{M_{A}}{M}\right) \delta\left(p^{2} - p'^{2}\right)$ $y \approx p'^{+} / P_{A}^{+}$ $S^{\alpha}(p) = \left\langle (A - 1)_{\alpha}, -\vec{p} \mid \hat{\psi}(0) \mid A \right\rangle$ $= 2\pi \delta(p_{0} - M - \varepsilon_{\alpha} + T_{R}) \left| \psi_{\alpha}(\vec{p}) \right|^{2}$

> Spectral function Quasi-elastic reaction A(e,e'p)A' $\rightarrow \varepsilon_{\alpha}$ Koltun sum rule: E/A = (T-e)/2 (2body force only)



The nucleon is scattered incoherently in case of

 $\ell_a \leq d \approx 2 \ \text{fm} \rightarrow x \geq 0.1$







Convolution form:

$$f_{a/A}(x) = \sum_{j,\alpha} \int dy dz \delta(x - yz) \int dp^2 D_{j/A}^{\alpha}(y, p^2) f_{a/j}(z, p^2)$$

Assumptions in the convolution model:

- on-mass shell approximation $\rightarrow p^2 = M^2 \rightarrow if$ the binding is weak, OK?
- impulse approximation final state interactions and interference terms are ignored

If OK, we get
$$F_2^A(x) = \sum_{j,\alpha} \int_x^A dy D_{j/A}^{\alpha}(y) F_2^{j}(x/y)$$

Model-dependent calculations:

- 1 Off-mass shell effect by Kulagin et al. \checkmark
- 2 Final state interactions (q-ex.) by Hoodhboy et al. \checkmark
- ③ "Off-mass" shell (↓) + final state interaction (MFA) by Saito et al. ↑



Ignored diagrams in the convolution



Note: Deuteron is also different from the average of proton and neutron — small EMC effect.

KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 10/19



Nonrelativistic calculation (by Li, Liu, Brown)

KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 11/19

Relativistic calculation (by Smith, Miller)





KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 12/19

3-2. Shadowing effect at small x



Shadowing region $\rightarrow \ell_c \geq d \approx 2 \ fm \rightarrow x \leq 0.1$

DIS occurs coherently: $F_2^A(x) \approx \sigma_{\gamma^*A} < A \times \sigma_{\gamma^*N} (\sigma_{\gamma^*A} \approx A^{0.8} \times \sigma_{\gamma^*N})$



 $|A_a / A_b| >> 1$ for x > 0.1 << 1 for x < 0.1

for small x, the photon is supposed to be converted into vector mesons VMD \rightarrow surface interaction $A^{2/3} \approx A^{0.8}$



KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 13/19

Shadowing effect (by Piller et al.)





KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 14/19

3-3. Anti-shadowing?



Anti-shadowing region $\rightarrow 0.1 \le x \le 0.2$

An enhancement at small x region \rightarrow pion field enhancement ???



Recent data of the giant Gamow-Teller states \rightarrow the Landau-Migdal parameters $g'_{NN} \approx 0.59$, $g'_{N\Delta} \approx 0.18 + 0.05 g'_{\Delta\Delta}$



KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 15/19

Parameterization of the EMC effect (by M. Hirai et al.)







KEK研究会『超高エネルギー宇宙線とハドロン構造2008』 16/19

4. Possible explanations — what do we need ?











5. Summary



- 1. The quark distribution in a nucleus is different from that in the free nucleon:
 - about 20% reduction at x ~ 0.7-0.8
 - at small x, the structure function is reduced due to shadowing
 - for large x, the EMC ratio is very enhanced because of Fermi motion and short-range correlations
- 2. The energy-momentum distribution of a nucleon in a nucleus is vital to explain the EMC effect, but its effect is insufficient ?
 the internal structure of a nucleon is modified in a nucleus ?
- 3. The sea quark is enhanced in a nucleus around x ~ 0.15 ?
 cf. the Drell-Yan result
- 4. At large x (>1), what happens ?

